The work forms given are essentially those in Ageton, plus about eight new ones. No warning is given of the difficulties which arise when the foot of the perpendicular lies near a pole; the smaller interval of the argument makes this less serious than in Ageton's tables.

The author's Case 7 (Latitude when the Sun or Star is near the Meridian) and Case 9 (Longitude when the Sun or Star is near the Prime Vertical) will probably appeal to many practical navigators. It is unfortunate that the author did not warn the reader of the troubles which arise in these cases if the altitude is near $90^{\circ}$.

The formula used in Case 9 is

$$
2 B(t / 2)=B(s)+B(s-z)-B(d)-B(L)
$$

where $t$ and $d$ are the meridian angle and declination respectively of the celestial body, $L$ is the observer's latitude, $z$ is $90^{\circ}-H$ or the body's zenith distance, $s=(z+d+L) / 2$. Table $S$, based upon this formula, gives $t$ and $z$ for Betelgeuse, each to the nearest minute of arc for $L, 0\left(1^{\circ}\right) 60^{\circ}$ and $H, 20^{\circ}\left(5^{\circ}\right) 35^{\circ}$. This table is quite similar to, and possesses many of the advantages of, Tafeln zur astronomischen Ortsbestimmung, by Arnold Kohlschütter, Berlin, 1913. A set of similar star tables prepared today for the northern temperate zone would make a valuable addition to navigational literature. They would possess the great advantage of a direct approach, allowing the navigator to enter a table with the observed altitude as an argument, and with an assumed latitude or longitude, find the corresponding longitude or latitude.

A brief examination of the principal table (II) indicates that the tabular values are much more reliable than those in Ageton.

Charles H. Smiley

518[V].-G. Moretti, "Scie piane turbolente," L'Aerotecnica, v. 27, 15 June 1947, p. 210-221. $20.5 \times 29 \mathrm{~cm}$.

The tables, p. 219-220, are (1) of $M(x \mid \alpha, \xi), N(x \mid \alpha, \xi), P(x \mid \alpha, \xi), \alpha=1.069, \xi=.765$, for $x=[0(.1) 2,2.2,2.5,3 ; 3-4 D]$, where
$M=\left[\alpha^{2} e^{-x^{2}} / C(\xi)\right]\left|\begin{array}{c}A(x) C(x) \\ A(\xi) C(\xi)\end{array}\right|, \quad N=\left[\alpha^{2} e^{-x^{2}} / C(\xi)\right]\left|\begin{array}{l}B(x) C(x) \\ B(\xi) C(\xi)\end{array}\right|$,
$P=\left[2 \alpha e^{-x^{2}} / C(\xi)\right]\left|\begin{array}{cc}x^{2}-\frac{1}{2} & C(x) \\ \xi^{2}-\frac{1}{2} & C(\xi)\end{array}\right|, \quad A(x)=\frac{1}{\frac{1}{2}} e^{-x^{2}}+x \int_{0}^{x} e^{-t^{2}} d t$,
$B(x)=\frac{1}{2} e^{-x^{2}}-x \int_{0}^{x} e^{-t^{2}} d t, \quad C(x)=\frac{1}{2} e^{x^{2}}-x \int_{0}^{x} e^{t^{2}} d t ;$
(2) of $y(x \mid \alpha, \xi, \mu)=M(x \mid \alpha, \xi)+\mu N(x \mid \alpha, \xi)$ and of $y(x \mid \alpha, \xi, \lambda)=M(x \mid \alpha, \xi)+\lambda P(x \mid \alpha, \xi)$, where $\mu=1.5369,1,0,-2,-5,-10, \lambda=-1.2155,-1(1) 2,5,10$ and in both cases $\alpha=1.069, \xi=.765, x=[0(.1) 2,2.2,2.5,3 ; 3 \mathrm{D}]$.

Extracts from text

## MATHEMATICAL TABLES-ERRATA

References have been made to Errata in RMT 477 (Cazzola), 484 (NBSCL), 486 (Anan'ev), 495 (Chebyshev, Whittaker \& Robinson), 503 (Magnus \& Oberhettinger), 506 (U. S. Navy), 510 (Skinner); N 87 (German Tables), 92 (Fontoura \& Penteado).
124. J. R. Airey, "Tables of the Bessel functions $J_{n}(x), " n=0(1) 13$, $x=[6.5(.5) 16 ; 10 \mathrm{D}]$, BAAS, Report, 1915, p. 30-32.

We note the following 51 cases, for $n=2(1) 13$, where Airey and Harvard (RMT 380, 440, 501) differ:

| $n$ | $x$ | Airey | Harvard |
| :---: | :---: | :---: | :---: |
| 2 | 6.5 | -0.30743 02368 | -0.30743 0390630. |
| 3 |  | -0.03534 65366 | -0.03534 $6631285 .$. |
| 4 |  | +0.27480 26645 | +0.27480 27310 |
| 5 |  | 0.3735652006 | 0.3735653771 |
| 6 |  | 0.2999130288 | 0.2999132338 |
| 7 |  | 0.1801203909 | 0.1801205930 |
| 8 |  | 0.0880385825 | 0.0880388126 |
| 9 |  | 0.0365899659 | 0.0365903304 |
| 10 |  | 0.0132874770 | 0.0132882562 |
| 11 |  | 0.0042945787 | 0.0042966118 |
| 12 |  | 0.0012480202 | 0.0012541220 |
| 13 |  | -0.00031 35057 | +0.00033 39927 |
| 4 | 8.5 | -0.20767 73541 | -0.20770 08835 |
| 5 |  | +0.06715 51647 | +0.06713 30194 |
| 6 |  | 0.2866834302 | 0.2866809063 |
| 7 |  | 0.3375743838 | 0.3375929660 |
| 8 |  | 0.2693214373 | 0.2693545671 |
| 9 |  | 0.1693836158 | 0.1694273956 |
| 10 |  | 0.0893732784 | 0.0894328589 |
| 11 |  | 0.0409064511 | 0.0410028606 |
| 12 |  | 0.0165022421 | 0.0166921921 |
| 13 |  | 0.0056881149 | 0.0061280346 |
| 11 | 9.5 | 0.0896964138 | 0.0896964137 |
| 12 |  | 0.0426916061 | 0.0426916060 |
| 13 |  | 0.0181560647 | 0.0181560646 |
| 6 | 11.5 | -0.24508 38970 | -0.24508 14040 |
| 7 |  | $-0.0846270668$ | -0.08462 44653 |
| 8 |  | +0.14205 96418 | +0.14206 03158 |
| 9 |  | 0.2822752641 | 0.2822736003 |
| 10 |  | 0.2997625107 | 0.2997592326 |
| 11 |  | 0.2390508414 | 0.2390468041 |
| 12 |  | 0.1575521425 | 0.1575476971 |
| 13 |  | 0.0897536298 | 0.0897483898 |
| 8 | 13.5 | -0.20367 10209 | -0.20367 08728 |
| 9 |  | -0.02727 91962 | -0.02727 93539 |
| 10 |  | +0.16729 87593 | +0.1672984008 |
| 11 |  | 0.2751292100 | 0.2751288367 |
| 12 |  | 0.2810599533 | 0.2810597034 |
| 13 |  | 0.2245329292 | 0.2245328582 |
| 9 | 14 | -0.11430 71982 | -0.11430 71981 |
| 3 | 14.5 | -0.21021 97923 | -0.21021 9792422. |
| 4 |  | -0.02612 20608 | -0.02612 25583 |
| 5 |  | +0.19580 76209 | +0.19580 73465 |
| 6 |  | 0.1611617993 | 0.1611621076 |
| 7 |  | -0.06243 23387 | -0.06243 18091 |
| 8 |  | -0.22144 12987 | -0.22144 10957 |
| 9 |  | -0.18191 66806 | -0.18191 69861 |
| 10 |  | -0.00438 63048 | -0.00438 68871 |
| 11 |  | +0.17586 66050 | +0.17586 61074 |
| 12 |  | 0.2712183952 | 0.2712182225 |
| 13 |  | 0.2730466008 | 0.2730468125 |

## R. C. A.

The discrepancies indicated above have all been tested by comparison with the original calculations, performed under the supervision of L. J. Comrie, for the forthcoming BAAS, Mathematical Tables, Bessel Functions, part II. In every case the Harvard value is confirmed, although incompletely in 2 cases. These two cases are:
(i) $J_{9}(14)$ where Comrie's 12 -decimal value ends ... 81 50. The value given by Meissel (see Gray, Mathews \& MacRobert A Treatise on Bessel Functions) has . . 81 497. .., so that Airey's error may presumably be explained as a rounding-off to 12 decimals, followed by another rounding-off to 10 decimals without reference back to Meissel's table, which Airey quotes as his source for integer $x$. This error is, however, quite trivial.
(ii) $J_{9}$ (14.5). The B.A. 12-decimal value ends ...61 $51 \ldots$, but is subject to a possible error of 2 units or so in the last figure. It has not seemed worthwhile to pursue the matter further.

It is disconcerting to find serious errors of this nature in any work of Airey's, even though it is comparatively early work. It is thus desirable to investigate more closely. A large pile of Airey's manuscripts were handed over to Dr. Comrie after Airey's death in 1937, but the calculations for this particular table do not seem to have been included.

The published values were therefore tested by formation of the values of

$$
E=x \bar{J}_{n-1}(x)-2 n \bar{J}_{n}(x)+x \bar{J}_{n+1}(x)
$$

in which $\bar{J}_{n}(x)$ is used to denote Airey's tabulated value.
The following values are seriously in error.

| $x$ | $n$ | $E$ in units of the 10th decimal |  |
| :---: | :---: | :---: | :---: |
| 6.5 | 1 | + | 10000 |
|  | 12 | $+$ | 638 |
| 8.5 | 3 | +20 | 00002 |
| 11.5 | 5 | - 2 | 86690 |
| 13.5 | 7 | - | 19996 |
|  | 8 | + | 44998 |
| 14.5 | 3 | $+$ | 72132 |

Apart from these the greatest value of $|E|$ is 18 units for $x=14.5, n=2$, arising from the small error at $n=3$.

Thus most of Airey's errors are readily explained. All of them, except the end figure ones, arise from the major computation errors indicated in the list. Four of these errors (the 1st, 3 rd, 5th and 6th) are almost certainly accounted for as errors in addition, subtraction or transcription. The second seems to fall into the same class, since $638=650-12=100 x$ -12 , and 12 units of the 10 th decimal is a residual that might reasonably arise from round-ing-off errors. The cause of the other two is, however, obscure.

It therefore appears probable that the errors have arisen through the publication of an early work without applying any check, since any check must surely have brought to light at least one of these errors, and so have directed Airey's attention to the need for a thorough examination. It is unfortunate that he did not carry the values for $x=6.5$ a little beyond $n=13$, for the error is within 4 or 5 lines of taking complete charge, and swamping the true value of $J_{n}(x)$ entirely.

J. C. P. Miller

## 1 January 1948

Editorial Note: There are 49 titles (1911-1938) in the list of published mathematical tables by J. R. Airey (1868-1937). The title here in question is no. 11. Dr. Miller's report on Errata in the 6D portion of this table will appear in MTAC 23.
125. Albert Gloden, "Table de factorisation des Nombres $X^{4}+1$ dans l'intervalle $1000<X \leq 3000^{\prime \prime}$; see RMT 348, MTAC, v. 2, p. 211.
Corrections of five errors in this table are as follows:
P. $731120^{4}+1=17 \times 89 \times 1039999$ 577,
p. $82 \quad 2310^{4}+1=17 \times 41 \times 40852171033$,
p. $76 \quad \frac{1}{2}\left(1417^{4}+1\right)=90841 \times 22190521$,
p. $77 \frac{1}{2}\left(1623^{4}+1\right)=17 \times 204077517313$,
p. $82 \quad \frac{1}{2}\left(2313^{4}+1\right)=433 \times 593 \times 55735249$.

17 of the blank spaces in the table may now be filled in as follows:
P. $74 \quad 1140^{4}+1=592649 \times 2849849$,
p. $74 \quad 1242^{4}+1=565921 \times 4204$ 657,
p. $75 \quad 1354^{4}+1=593081 \times 5667$ 097,
p. $761444^{4}+1=539089 \times 8065073$,
p. $79 \quad 1904^{4}+1=595201 \times 22080257$,
p. $82 \quad 2190^{4}+1=530713 \times 43342777$,
p. $83 \quad 2332^{4}+1=518417 \times 57047281$,
p. $83 \quad 2406^{4}+1=17 \times 511457 \times 3854113$,
p. $74 \quad \frac{1}{2}\left(1229^{4}+1\right)=597769 \times 1908279$,
p. $75 \quad \frac{1}{2}\left(1299^{4}+1\right)=527377 \times 2699513$,
p. $79 \quad \frac{1}{2}\left(1867^{4}+1\right)=563081 \times 10788881$,
p. $79 \quad \frac{1}{2}\left(1869^{4}+1\right)=522553 \times 11675537$,
p. $80 \frac{1}{2}\left(2005^{4}+1\right)=2089 \times 3868023217$,
p. $80 \frac{1}{2}\left(2055^{4}+1\right)=17 \times 572233 \times 916633$,
p. $82 \frac{1}{2}\left(2189^{4}+1\right)=526249 \times 21815329$,
p. $87 \frac{1}{2}\left(2895^{4}+1\right)=570001 \times 61615313$,
p. $87 \quad \frac{1}{2}\left(2969^{4}+1\right)=520529 \times 74639009$.

Furthermore, besides the factors given at the following 17 entries, the remaining factor in each case is a 12-figure prime of the form $8 k+1<600000^{2}$ : (a) $X^{4}+1, \mathrm{p} .77, X=1562$; p. 79, $X=1818$; p. $79, X=1848$; p. $82, X=2262$; p. $82, X=2302$; p. $84, X=2468$; p. 84, $X=2476 ;$ p. 86, $X=2808 ;$ p. $88, X=3000$. (b) $\frac{1}{2}\left(X^{4}+1\right)$, p. 78, $X=1709 ;$ p. 78, $X=1715 ;$ p. 82, $X=2211 ;$ p. 82, $X=2299 ;$ p. $84, X=2533 ;$ p. $84, X=2577$; p. 85 , $X=2669$; p. 85, $X=2683$.

## A. Gloden

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126. O. A. Walther, "Bemerkungen über das Tschebyscheffsche Verfahren zur numerischen Integration," Skandinavisk Aktuarietidskrift, v. 13, 1930, p. 168-192.
On p. 177-179 are given the roots of Chebyshev's polynomials of degree $n$ (those employed in numerical integration with equal weight factors) for $n=[1(1) 7 ; 10 \mathrm{D}], n=[8(1)$ $10 ; 5 \mathrm{D}]$. This fact was unknown to the writer when he also gave, among other quantities, the roots, to 10D, for those polynomials having only real roots (see RMT 495). The writer had relied on FMR, Index, p. 360, where no mention was made of Walther's roots which supersede the calculations of most of the authors cited there; hence the writer's statement on p . 192-193 regarding other tables of roots must be slightly modified in order to take Walther's tables into account.

Comparison of Walther's roots with the writer's, where they overlapped, revealed one appreciable error in Walther's calculations which was greater than a unit in the last place. On p. 178, for Walther's $n=4$ (corresponding to the writer's $n=5$ ) a pair of roots is given as $\pm 0.8324974841$, whereas it should have been $\pm 0.8324974870$.

Herbert E. Salzer

127. Dov Yarden, (a) "Table of Fibonacci numbers"; (b) "Table of the ranks of apparition in Fibonacci's sequence," Riveon Lematematika, v. 1, 1946, p. 35-37; 54 ; v. 2, Sept. 1947, p. 22. The errata below supplement those listed in MTAC, v. 2, p. 343-344.
(a) In the last number (5) of Riv. Lem., v. 1, June 1947, p. 99, are the following corrections in factorizations of $U_{n}$ and $V_{n}$ :

|  | $U_{n}$ | $n$ |
| :--- | :--- | :---: |
| $5 \cdot 28657 \cdot(3372041404278257761)$ | 483162952612010163284885 | 115 |
| $353 \cdot 709 \cdot 8969 \cdot 336419 \cdot 2710260697$ | 2046711111473984623691759 | 118 |


| $3 \cdot 347 \cdot 1270083883$ | $\boldsymbol{V}_{n}$ | $n$ |
| :--- | ---: | ---: |
| $2^{2} \cdot 19 \cdot 199 \cdot 991 \cdot 2179 \cdot 9901 \cdot 1513909$ | 1322157322203 | 58 |
| $2 \cdot 3^{2} \cdot 227 \cdot 29134601 \cdot(5608975608563)$ | 489526700523968661124 | 99 |
| $2^{2} \cdot 19 \cdot 79 \cdot 521 \cdot 859 \cdot(1052645985555841)$ | 667714778405043259651218 | 114 |
|  | 2828485190904971853895196 | 117 |

(b) On p. 22 of v. 2, line 13, the author noted the following corrections:
$p=1031$, for $2 \cdot 5 \cdot 103 \quad 1030$, read $2 \cdot 5 \cdot 103 \quad 206$;
$p=1231$, for $2 \cdot 3 \cdot 5 \cdot 41 \quad 1230$, read $2 \cdot 3 \cdot 5 \cdot 41 \quad 410$.

## UNPUBLISHED MATHEMATICAL TABLES

Reference has also been made to Unpublished Tables in RMT 485 (Glaisher), 491 (Gloden); Q24 (Wrench).
67[F].-P. Poulet, "Suites de totalics au depart de $n \leq 2000$." Hectographed copy on one side of each of 56 leaves, in possession of D. H. L. $20 \times 24.8 \mathrm{~cm}$.
By a "totalic series" or "aliquot series" is meant a sequence of positive integers, each term of which is the sum of the proper divisors of its predecessor. Two simple examples are 18, 21, 11, 1 1420, 1604, 1210, 1184, 1210, 1184, $\cdots$.
The first of these terminates with its fourth term; the second ultimately becomes periodic of period two. It has been conjectured ${ }^{1}$ that aliquot series either terminate or become periodic. The present tables show this to be the case for all such sequences whose "leaders" (first terms) do not exceed 2000, with the possible exception of about 25 series which are left unfinished. For each such leader are given those terms of the sequence which are $\leq n$. When a term $n_{1}$ finally falls below $n$, the reader is referred to the previous series whose leader is $n_{1}$. When the leader is a term of a previous series, reference is made to the leader of this series. Prime leaders are of no interest and are omitted. Beyond $n=200$ only abundant leaders $n$ are listed. Other leaders would have given second terms not greater than the leaders.

Some leaders generate unusually long sequences. The longest completed series is

$$
936,1794,2238,2250, \cdots, 74,40,50,43,1
$$

and runs to 189 terms, the largest term being

$$
33289162099526=2 \cdot 25943 \cdot 641582741
$$

Thus the three dots of this series represent a formidable calculation. It is due to B. H. Brown who (since 1940) also contributed many terms to several of the other still incomplete series. The incomplete series with the smallest leader is 276, 396, 696, 1104, $\cdots, 5641400009252, \cdots$ ( 58 terms).
Besides giving the terms in their decimal representation, the author gives their canonical factorization into primes. This table is an extension of a previous table of Dickson ${ }^{2}$ for leaders $\leq 1000$.

D. H. L.

[^0]
[^0]:    ${ }^{1}$ L. E. Drckson, History of the Theory of Numbers, v. 1, Washington, Carnegie Institution, 1919 ; offset print, New York, Stechert, 1934, p. 48-49.
    ${ }^{2}$ L. E. Dickson, "Theorems and tables on the sum of the divisors of a number," Quart. Jn. Math., v. 44, 1913, p. 267-272. For additions and corrections see P. Poulet, La Chasse aux nombres, Brussels, v. 1, 1929, p. 69-72; v. 2, 1934, p. 187-8.

    68[G].-Herbert E. Salzer, Chebyshev Polynomials, ms. in possession of the author at NBSCL.
    C. Lanczos, in his "Trigonometric interpolation of empirical and analytical functions," Jn. Math. Phys., v. 17, 1938, p. 140, gave the coefficients of the Chebyshev polynomials $C_{n}(x)$ adjusted to the range [0,1], up to $n=10$. Due to their importance, these coefficients

