A SIMPLE APPROACH TO THE SUMMATION OF CERTAIN SLOWLY CONVERGENT SERIES

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ABSTRACT. Summation of series of the form $\sum_{k=1}^{\infty} k^{\nu-1} r(k)$ is considered, where $0 \leq \nu \leq 1$ and r is a rational function. By an application of the Euler-Maclaurin summation formula, the problem is reduced to the evaluation of Gauss' hypergeometric function. Examples are given.

1. INTRODUCTION

In a recent paper [3], Gautschi has considered series of the type

$$(1.1) \qquad \sum_{k=1}^{\infty} k^{\nu-1} r(k),$$

where $0 < \nu \leq 1$ and r is a rational function

$$r(k) = \frac{p(k)}{q(k)},$$

p, q being real polynomials, deg $p \leq$ deg q. It is assumed that the zeros of q all have nonpositive real parts. By obtaining the fraction decomposition of r, the problem can be simplified to considering rational functions of the form

$$r(k) = \frac{1}{(k+a)^m} \qquad (\Re a \geq 0,\ m \geq 1).$$

The fractional power $k^{\nu-1}$ in (1.1) may be generalized to $(k+b)^{\nu-1}$, $\Re b \geq 0$; we thus consider series of the type

$$(1.2) \qquad \sum_{k=1}^{\infty} \frac{(k+b)^{\nu-1}}{(k+a)^m} \qquad (m \geq 1,\ 0 < \nu < 1).$$

In the cited paper, the sum of the series (1.2) is expressed as a weighted integral over $\mathbb{R}_+$ of certain special functions related to the incomplete gamma function. Gaussian quadrature is applied to the integral, using $w_\nu(t) = t^{-\nu} \varepsilon(t)$, where $\varepsilon(t) = t(e^t - 1)^{-1}$ is the Einstein function, as a weight function on $[0, \infty)$. Convergence of the quadrature formula can be shown. Nevertheless,

the application of the method is not a simple task, in general. The main difficulty is connected with the evaluation of the incomplete gamma function. For $\nu = \frac{1}{2}$, this function is expressible in terms of Dawson's integral F, which can be computed to high accuracy.

In this paper, we show that the problem of summing the series (1.2) may be solved using very simple tools. Application of the Euler-Maclaurin formula reduces the problem to the evaluation of Gauss' hypergeometric function $_2F_1$.

2. RESULTS

Let s denote the sum of the series (1.2). We can write

$$(2.1) \qquad s = \sum_{k=1}^{n-1} \frac{(k+b)^{\nu-1}}{(k+a)^m} + r_n,$$

where

$$(2.2) \qquad r_n := \sum_{k=n}^{\infty} \frac{(k+b)^{\nu-1}}{(k+a)^m}.$$

Application of the Euler-Maclaurin formula (see, e.g., [1, 23.1.30]) to (2.2) gives

$$(2.3) \qquad r_n - \int_n^{\infty} f(x)\,dx - \frac{1}{2}f(n) \sim -\sum_{k=1}^{\infty} \frac{B_{2k}}{(2k)!} f^{(2k-1)}(n),$$

where

$$f(x) := \frac{(x+b)^{\nu-1}}{(x+a)^m},$$

and B_j is the jth Bernoulli number. Using a formula of [4, 3.194.2], we obtain

$$(2.4) \qquad \int_n^{\infty} f(x)\,dx = \frac{1}{m-\nu}(n+b)^{\nu-m}\,_2F_1\left(\begin{matrix} m,\ m-\nu \\ m+1-\nu \end{matrix} \middle| \frac{b-a}{n+b}\right).$$

Equations (2.3), (2.4) and the formula

$$f^{(r)}(x) = \frac{(-1)^r(1-\nu)_r}{(x+a)^m(x+b)^{r+1-\nu}} \sum_{j=0}^{r}(-1)^j\binom{r}{j}\frac{(m)_j}{(\nu-r)_j}\left(\frac{x+b}{x+a}\right)^j \qquad (r \geq 0)$$

imply that

$$(2.5) \qquad r_n = r_n^{(q)} + O(n^{-2q-1}) \qquad (q \geq 1),$$

where

$$(2.6) \qquad \begin{aligned} r_n^{(q)} &= \frac{(n+b)^{\nu-m}}{m-\nu}\,_2F_1\left(\begin{matrix} m,\ m-\nu \\ m+1-\nu \end{matrix} \middle| \frac{b-a}{n+b}\right) \\ &\quad + \frac{(n+b)^{1-\nu}}{(n+a)^m}\left\{\frac{1}{2} + \sum_{k=1}^{q}\frac{B_{2k}(1-\nu)_{2k-1}}{(2k)!(n+b)^{2k-1}}\sigma_k\right\}. \end{aligned}$$

Here,

$$\sigma_k := \sum_{j=0}^{2k-1}(-1)^j\binom{2k-1}{j}\frac{(m)_j}{(\nu-2k+1)_j}\left(\frac{n+b}{n+a}\right)^j \qquad (k=1,2,\ldots),$$

and $(c)_j := c(c + 1) \cdots (c + j - 1)$ is Pochhammer's symbol.

The function $_2F_1$ appearing in the equation (2.6) can be evaluated by applying some standard linear transformations (see [1, 15.3.4, 15.3.7]) which brings this function into a form having a fast convergent power series expansion.

3. EXAMPLES

We made several tests, taking various combinations of the parameters. In all examples below we take $q = 4$ (cf. (2.5)), i.e., we compute the approximation

$$(3.1) \qquad \tilde{s}_n := \sum_{k=1}^{n-1} \frac{(k + b)^{\nu - 1}}{(k + a)^m} + r_n^{(4)}$$

to the sum s of the series (1.2), where $r_n^{(4)}$ is defined by (2.6).

All the computations were done on a personal 486-based computer, in extended arithmetic (about 18 decimal places).

Example 3.1. Let $a > 0$, $b = 0$, $\nu = \frac{1}{2}$. This is the case discussed in [3, §3]. Note that the hypergeometric function $_2F_1(z)$ in the formula (2.6) simplifies in this case to the arctan function. The results for $m = 1$ are shown in Table 3.1. Convergence is very fast, even for large a, in which case Gautschi's method of [3] needs some extra effort to achieve good accuracy. This remains true for other values of $\nu \in (0, 1)$ and $b > 0$; in Table 3.2 we give a selection of results obtained for $\nu = \frac{9}{10}$, $b = \frac{1}{2}$ and $m = 1$.

A similar rate of convergence is observed for $m > 1$.

TABLE 3.1. Approximations (3.1) to the sum of (1.2) for $a = 1$, $8, 64, b = 0, \nu = \frac{1}{2}, m = 1$

n	$a = 1$	$a = 8$	$a = 64$
5	1.8600250788	0.93137293396	0.3699316982450
10	1.8600250792207	0.93137293400304	0.369931698249664
15	1.860025079221182	0.9313729340031025	0.3699316982496709
20	1.8600250792211898	0.93137293400310378	0.36993169824967112
30	1.86002507922119030	0.93137293400310387	0.36993169824967113
40	1.86002507922119031	0.93137293400310387	

TABLE 3.2. Approximations (3.1) to the sum of (1.2) for $a = 1$, $8, 64, b = \frac{1}{2}, \nu = \frac{9}{10}, m = 1$

n	$a = 1$	$a = 8$	$a = 64$
5	9.61542462126	8.1705856754462	6.6951505182819
10	9.61542462140421	8.170585675449024	6.6951505182822286
15	9.6154246214045207	8.17058567544903510	6.69515051828222948
20	9.61542462140452727	8.17058567544903549	6.69515051828222951
30	9.61542462140452771	8.17058567544903553	6.69515051828222951
40	9.61542462140452772	8.17058567544903553	

Example 3.2. The method works well also for complex a. For $a = i\alpha$, $\alpha > 0$, $b = 0$, $\nu = \frac{1}{2}$, $m = 1$, we obtained the results shown in Table 3.3. For each n, the first entry is $\Re\tilde{s}_n$, the second $\Im\tilde{s}_n$ (cf. (3.1)). The nature of the convergence is the same as in the preceding examples. In contrast with the method of [3],

TABLE 3.3. Approximations (3.1) to the sum of (1.2) for $a = i\alpha$, $\alpha > 0$, $b = 0$, $\nu = \frac{1}{2}$, $m = 1$

n	$\alpha = 1$	$\alpha = 8$	$\alpha = 64$
5	2.0061526550	0.782147849843	0.27762942965433
	−0.796488122	−0.6029037623	−0.254862241652
10	2.006152655226	0.78214784984205	0.27762942965430958
	−0.7964881235693	−0.6029037624090	−0.25486224165720
15	2.00615265522740	0.782147849842075	0.27762942965430952
	−0.796488123569842	−0.60290376240912464	−0.2548622416572145
20	2.006152655227413	0.782147849842075	0.27762942965430952
	−0.796488123569847	−0.60290376240912460	−0.25486224165721467
25	2.00615265522741422	0.78214784984207493	0.27762942965430952
	−0.79648812356984801	−0.60290376240912467	−0.25486224165721468
30	2.00615265522741426	0.78214784984207492	
	−0.79648812356984802	−0.60290376240912468	

TABLE 3.4. Approximations (3.1) to the sum of (1.2) for $a = 1 + b + \omega i$, $\omega > 0$, $b = \frac{1}{2}$, $\nu = \frac{9}{10}$, $m = 1$

n	$\omega = 1$	$\omega = 8$	$\omega = 64$
5	9.27126436493	8.1408914758952	6.625908132988981
	−0.3966333402	−1.05441141121	−1.0214791502713
10	9.2712643649515	8.14089147589500	6.62590813298898329
	−0.3966333403072	−1.05441141122257	−1.021479150271591
15	9.271264364951661	8.140891475895024	6.62590813298898329
	−0.396633340307402	−1.05441141122257	−1.02147915027159224
20	9.2712643649516653	8.140891475895024	6.62590813298898329
	−0.3966333403074054	−1.0544114112225691	−1.02147915027159226
25	9.27126436495166560	8.14089147589502396	
	−0.39663334030740557	−1.05441141122256908	
30	9.27126436495166563	8.14089147589502394	
	−0.39663334030740558	−1.05441141122256909	

no difficulties are observed in case of large α. The same accuracy is obtained for other values of a, b, ν, and m; Table 3.4 gives the results obtained by the proposed method for $a = 1 + b + \omega i$, $\omega > 0$, $b = \frac{1}{2}$, $\nu = \frac{9}{10}$, and $m = 1$.

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