

## CORRECTION TO "TOTALLY GEODESIC EINSTEIN SPACES"<sup>1</sup>

AARON FIALKOW

The coordinate system (p. 427) in which  $H = x^n$  for some fixed value of  $y$  and  $f_{n\lambda} = 0$  exists if and only if  $f^{ij}H_{,i}H_{,j} \neq 0$  for this value of  $y$ . Hence Theorem 3.1 is valid only if this inequality holds. The remaining case, namely,

$$(3.15) \quad f^{ij}H_{,i}H_{,j} = 0$$

for every  $y$  can arise only if  $c = 0$ , as may be seen by differentiating (3.15) covariantly with respect to  $k$  and using (3.7). We note that, in accordance with (3.8) and (3.9),  $c = 0$  implies that  $a = b = 0$ . To obtain the analogue of Theorem 3.1 for the case in which (3.15) holds, we proceed in a manner analogous to that in H. W. Brinkmann, loc. cit., pp. 131–135 or A. Fialkow, *Conformal geodesics*, Transactions of this Society, vol. 45 (1939), p. 473. By these methods, we find a coordinate system such that  $H = x^n$  for a fixed value of  $y$  and

$$\begin{aligned} f^{ns} &= 0, & f^{nn} &= 0, & f^{(n-1)n} &= 1, \\ f_{i(n-1)} &= 0, & f_{(n-1)(n-1)} &= 0, & f_{(n-1)n} &= 1, \end{aligned}$$

where  $s, t = 1, 2, \dots, n-2$ . In this coordinate system, the characteristic condition (3.7) becomes  $\partial g_{ij}/\partial x^{n-1} = 0$ . (In the Transactions paper, this last equation appears incorrectly as  $\partial g_{st}/\partial x^{n-1} = 0$ .)

If the  $f_{ij}$  are to be the components of the metric tensor of an Einstein space  $E_n$ , then, as was shown by Brinkmann, the first fundamental form of  $E_n$  may be written as

$$(3.16) \quad \begin{aligned} f_{st} &= h_{st}(x^s, x^n), & f_{sn} &= 0, & f_{nn} &= 0, \\ f_{(n-1)n} &= 1, & f_{s(n-1)} &= 0, & f_{(n-1)(n-1)} &= 0, \end{aligned}$$

where  $h_{st}dx^sdx^t$  with  $x^n$  constant is the first fundamental form of an Einstein space  $E_{n-2}$  of zero mean curvature, and the components of the tensor  $h_{st}$  satisfy certain partial differential equations. According to Brinkmann, the conditions (3.16) are the necessary and sufficient conditions that  $E_n$  be conformal to another Einstein space by means of a transformation  $d\bar{s} = \sigma ds$  with  $\Delta_1\sigma = f^{ij}\sigma_{,i}\sigma_{,j} = 0$ . We note that the most general solution for  $H$  of the form  $H = H(x^n, y)$  is given by (3.13). Now this solution  $H(x^n, y)$  must involve  $x^n$  by the hypothesis

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of §3 and must also depend upon  $y$  since  $g^{\alpha\beta}H_{,\alpha}H_{,\beta} \neq 0$  and, according to (3.1) and (3.15),  $g^{\alpha\beta}H_{,\alpha}H_{,\beta} = g^{(n+1)(n+1)}(\partial H/\partial y)^2$ . Hence the first fundamental form of the  $E_{n+1}$  which contains the  $\infty^1$  isometric  $E_n$ 's is

$$ds^2 = f_{ij}dx^i dx^j + e[\alpha(y)x^n + \beta(y)]^2 dx^{n+1}^2$$

where the  $f_{ij}$  satisfy (3.16) and  $\alpha \neq 0$  and  $(d\alpha/dy)^2 + (d\beta/dy)^2 \neq 0$ . These remarks show that Theorem 3.1 must be modified in the case where (3.15) holds so that the phrase " $\Delta_1\sigma \neq 0$ " is replaced by the phrase " $\Delta_1\sigma = 0$ ." Both cases may be included in the theorem:

**THEOREM.** *A one-parameter family of isometric  $E_n$ 's may be imbedded as  $\infty^1$  nonparallel, totally geodesic hypersurfaces of an  $E_{n+1}$  if and only if each  $E_n$  may be mapped conformally on another Einstein space. If  $a$  and  $b$  are the mean curvatures of  $E_{n+1}$  and  $E_n$  respectively, then  $nb = (n-1)a$ .*

BROOKLYN COLLEGE