

Miniaturized Tables of Bessel Functions*

By Yudell L. Luke

Abstract. In this report, we discuss the representation of bivariate functions in double series of Chebyshev polynomials. For an application, we tabulate coefficients which are accurate to 20 decimals for the evaluation of $(2z/\pi)^{1/2}e^z K_\nu(z)$ for all $z \geq 5$ and all ν , $0 \leq \nu \leq 1$. Since $K_\nu(z)$ is an even function in ν and satisfies a three-term recurrence formula in ν which is stable when used in the forward direction, we can readily evaluate $K_\nu(z)$ for all $z \geq 5$ and all $\nu \geq 0$. Only 205 coefficients are required to achieve an accuracy of about 20 decimals for the z and ν ranges described. Extension of these ideas for the evaluation of all Bessel functions and other important bivariate functions is under way.

1. Introduction. One of the difficulties associated with the development of tables of multiparameter functions is the enormously large number of entries needed to cover the ranges required in these parameters. Even if such values are prepared, their direct use on high-speed computers is limited because of storage, the necessity of table look up and interpolation. From the vantage point of automatic computation, the need is for efficient algorithms and small sets of coefficients to compute entries as they are needed. In the case of a two-parameter function, we can expand $f(z, \nu)$, for example, in the form $\sum C_n(\nu)P_n(z)$ where the $P_n(z)$'s are orthogonal polynomials in z and the $C_n(\nu)$'s depend only on ν . This has the advantage that tabulation of a two-parameter function is made to depend upon the tabulation of two other functions, each of a single parameter. Evaluation of sums involving orthogonal polynomials by backward recurrence is easy, provided the three-term recurrence formula for the orthogonal polynomials is known. This is certainly the case for the classical orthogonal polynomials. For many functions, computation of the $C_n(\nu)$'s for a given ν can also be accomplished by use of backward recurrence techniques as is evidenced by my work on the special functions [1]. We now propose to expand $C_n(\nu)$ in series of orthogonal polynomials. In this manner, $f(z, \nu)$ is expanded in a double series, and by giving a rather small set of coefficients, we can compress previous extensive tabulations of two-parameter functions onto a few pages.

In this report, the above ideas are applied for the evaluation of the modified Bessel function $K_\nu(z)$ in a double series of Chebyshev polynomials valid for $z \geq 5$ and $0 \leq \nu \leq 1$. Further work along these lines is continuing, so that, eventually, we will have coefficients for all the Bessel functions over extensive z and ν values. We will also apply these techniques to other important two-parameter functions. Extension of these concepts to three and more parameter functions is also contemplated.

Received June 14, 1970.

AMS 1969 subject classifications. Primary 6520, 6525, 3325; Secondary 4130.

Key words and phrases. Bessel functions, approximation of bivariate functions, approximation in series of Chebyshev polynomials.

* This work was supported by the United States Atomic Energy Commission under contract number AT(11-1)1619 with the Midwest Research Institute.

Copyright © 1971, American Mathematical Society

2. Chebyshev Expansions for $K_\nu(z)$. From the work in [1], we have

$$(1) \quad K_\nu(z) = (\pi/2z)^{1/2} e^{-z} \sum_{k=0}^{\infty} C_k(\nu, \lambda) T_k^*(\lambda/z), \quad \lambda \text{ fixed, } \lambda/z \leq 1, |\arg z| < 3\pi/2,$$

where $K_\nu(z)$ is the modified Bessel function of order ν and $T_k^*(x)$ is the "shifted" Chebyshev polynomial of the first kind. The coefficients $C_k(\nu, \lambda)$ satisfy the recurrence formula

$$(2) \quad \frac{2C_k(\nu, \lambda)}{\epsilon_k} = 2(k+1) \left\{ 1 - \frac{(2k+3)(k+3/2+\nu)(k+3/2-\nu)}{2(k+2)(k+1/2+\nu)(k+1/2-\nu)} - \frac{4\lambda}{(k+1/2+\nu)(k+1/2-\nu)} \right\} C_{k+1}(\nu, \lambda) \\ + \left\{ 1 - \frac{2(k+1)(2k+3-4\lambda)}{(k+1/2+\nu)(k+1/2-\nu)} \right\} C_{k+2}(\nu, \lambda) \\ - \frac{(k+1)(k+5/2+\nu)(k+5/2-\nu)}{(k+2)(k+1/2+\nu)(k+1/2-\nu)} C_{k+3}(\nu, \lambda), \\ \epsilon_0 = 1, \quad \epsilon_k = 2, \quad k > 0.$$

As the expansion formula (1) converges, for λ and ν fixed,

$$(3) \quad \lim_{k \rightarrow \infty} C_k(\nu, \lambda) = 0.$$

Further, since

$$(4) \quad \lim_{z \rightarrow \infty} (2z/\pi)^{1/2} e^z K_\nu(z) = 1,$$

we have

$$(5) \quad \sum_{k=0}^{\infty} (-1)^k C_k(\nu, \lambda) = 1.$$

In the reference cited, it is shown that if $|\arg \lambda| < \pi$, the coefficients $C_k(\nu, \lambda)$ are readily evolved by use of (2) in the backward direction. This is very efficient. Further, no prior values of $K_\nu(z)$ need be known. To aid in the application of (2), we have the asymptotic formula [2],

$$(6) \quad C_k(\nu, \lambda) = \frac{4(-1)^k \pi^{1/2} k^{-2/3} (2\lambda)^{1/6}}{3^{1/2} \Gamma(\frac{1}{2} + \nu) \Gamma(\frac{1}{2} - \nu)} \exp(-3k^{2/3} \{2\lambda\}^{1/3}) [1 + O(k^{-1/3})].$$

Next, we consider the expansion of the coefficients $C_k(\nu, \lambda)$ in series of Chebyshev polynomials. To this end, it is easier to consider a more general situation. Suppose that $f(x)$ is continuous in $[-1, 1]$ except for a finite number of bounded jumps. Then, $f(x)$ can be expanded in a convergent series as

$$(7) \quad f(x) = \frac{1}{2} c_0 + \sum_{r=1}^{\infty} c_r T_r(x), \quad -1 \leq x \leq 1, \\ c_r = (2/\pi) \int_{-1}^1 (1-x^2)^{-1/2} f(x) T_r(x) dx.$$

Here, $T_r(x)$ is the "unshifted" Chebyshev polynomial of the first kind. Simple and easy to evaluate forms for c_k are not often available. If $f(x) = (2z/\pi)^{1/2} e^z K_\nu(z)$,

$z(x + 1) = 2\lambda$, the situation is considerably simplified, in view of the recurrence formula (2) and the normalization relation (5). For the case when $f(x)$ is $C_k(\nu, \lambda)$, no recurrence formula is known, and if the integral representation in (7) is used, we must resort to numerical integration. Now, the Chebyshev polynomials satisfy two orthogonality relations with respect to summation. Use of expansion formulas based on these conditions is equivalent to the evaluation of (7) by trapezoidal-type numerical integration formulas. Indeed, we have the following results. If

$$(8) \quad f(x) = f_n(x) + R_n(x),$$

$$(9) \quad f_n(x) = \frac{1}{2}d_0 + \sum_{r=1}^{n-1} d_r T_r(x),$$

then

$$(10) \quad d_r = (2/n) \sum_{\alpha=0}^{n-1} f(x_\alpha) T_r(x_\alpha), \quad x_\alpha = \cos \theta_\alpha, \quad \theta_\alpha = (\pi/2n)(2\alpha + 1),$$

and

$$(11) \quad d_r = c_r + \sum_{s=1}^{\infty} (-1)^s \{c_{2sn-r} + c_{2sn+r}\}, \quad r = 1, 2, \dots, n - 1.$$

Again, if

$$(12) \quad f(x) = f_n(x) + S_n(x),$$

$$(13) \quad f_n(x) = \frac{1}{2}e_0 + \sum_{r=1}^{n-1} e_r T_r(x) + \frac{1}{2}T_n(x),$$

then

$$(14) \quad e_r = \frac{2}{n} \left[\frac{1}{2} \{f(1) + (-1)^r f(-1)\} + \sum_{\alpha=1}^{n-1} f(x_\alpha) T_r(x_\alpha) \right],$$

$$x_\alpha = \cos \varphi_\alpha, \quad \varphi_\alpha = \alpha\pi/n,$$

and

$$(15) \quad e_r = c_r + \sum_{s=1}^{\infty} \{c_{2sn-r} + c_{2sn+r}\}, \quad r = 0, 1, \dots, n - 1,$$

$$e_n = 2c_n + 2 \sum_{s=1}^{\infty} c_{(2s+1)n}.$$

Thus, d_r and e_r are approximations to c_r , and from (11) and (15)

$$(16) \quad c_r = \frac{1}{2}(d_r + e_r) + \sum_{s=1}^{\infty} \{c_{4sn-r} + c_{4sn+r}\}, \quad r = 0, 1, \dots, n - 1.$$

Proof of (8)–(16) is given in [3]. There, we also develop some general information concerning the asymptotic behavior of c_r for large r . In most cases of interest, (7) converges quite rapidly, so that for n sufficiently large, c_r is well approximated by d_r or e_r , and an improved approximation is simply the arithmetic mean of the latter two quantities.

For application of (8)–(16) to $C_k(\nu, \lambda)$ valid for $0 \leq \nu \leq 1$, we write

$$(17) \quad C_k(\nu, \lambda) = \sum_{r=0}^{\infty} L_{r,k}(\lambda) T_r^*(\nu), \quad 0 \leq \nu \leq 1.$$

Coefficients in the Expansion of

$$K_{\nu}(z) = (\pi/2z)^{\frac{1}{2}} e^{-z} \sum_{k=0}^{\infty} C_k(\nu) T_k^*(5/z), \quad z \geq 5, \quad ,$$

$$C_k(\nu) = \sum_{r=0}^{\infty} L_{r,k} T_r^*(\nu), \quad 0 \leq \nu \leq 1. \quad .$$

r	$L_{r,k}, k = 0$				r	$L_{r,k}, k = 1$			
0	1.00607	66859	78294	33189	0	0.00602	15036	04662	74176
1	0.02367	20980	36879	87423	1	0.02325	81816	55108	78143
2	0.00608	97384	81944	49510	2	0.00603	86674	65067	42659
3	0.00009	90242	93398	39463	3	0.00012	93931	45178	04993
4	0.00001	30722	50003	18810	4	0.00001	71930	09049	19380
5	0.00000	02196	13551	61367	5	0.00000	03226	32525	19147
6	0.00000	00197	68831	32376	6	0.00000	00291	84000	17904
7	0.00000	00003	26575	97550	7	0.00000	00005	11729	27069
8	0.00000	00000	22498	49793	8	0.00000	00000	35388	74773
9	0.00000	00000	00360	09504	9	0.00000	00000	00587	87945
10	0.00000	00000	00020	21390	10	0.00000	00000	00033	10263
11	0.00000	00000	00000	31189	11	0.00000	00000	00000	52390
12	0.00000	00000	00000	01484	12	0.00000	00000	00000	02499
13	0.00000	00000	00000	00022	13	0.00000	00000	00000	00038
14	0.00000	00000	00000	00001	14	0.00000	00000	00000	00002

r	$L_{r,k}, k = 2$				r	$L_{r,k}, k = 3$			
0	-0.00005	37236	98553	04638	0	0.00000	13961	41734	51613
1	-0.00039	80431	75117	81456	1	0.00001	49927	12492	36296
2	-0.00004	97465	49910	04795	2	0.00000	12681	70364	79510
3	0.00002	89205	14415	32315	3	-0.00000	13610	18292	01200
4	0.00000	39880	78060	13616	4	-0.00000	01265	51139	39276
5	0.00000	01186	02610	05376	5	0.00000	00141	24509	01269
6	0.00000	00109	46041	18670	6	0.00000	00014	24282	69365
7	0.00000	00002	37314	19859	7	0.00000	00000	55693	42227
8	0.00000	00000	16641	85427	8	0.00000	00000	04058	21783
9	0.00000	00000	00313	12644	9	0.00000	00000	00100	45462
10	0.00000	00000	00017	82002	10	0.00000	00000	00005	85827
11	0.00000	00000	00000	30624	11	0.00000	00000	00000	11875
12	0.00000	00000	00000	01473	12	0.00000	00000	00000	00581
13	0.00000	00000	00000	00024	13	0.00000	00000	00000	00010
14	0.00000	00000	00000	00001					

r	$L_{r,k}, k = 4$				r	$L_{r,k}, k = 5$			
0	-0.00000	00589	09792	34012	0	0.00000	00033	63643	06632
1	-0.00000	08184	31644	38871	1	0.00000	00568	52171	72673
2	-0.00000	00530	50266	39835	2	0.00000	00030	14215	93849
3	0.00000	00808	85729	57198	3	-0.00000	00058	78142	93966
4	0.00000	00057	60293	40591	4	-0.00000	00003	42384	83406
5	-0.00000	00013	30534	35272	5	0.00000	00001	16340	34571
6	-0.00000	00000	98969	45233	6	0.00000	00000	07004	16824
7	0.00000	00000	02855	70119	7	-0.00000	00000	00591	12471
8	0.00000	00000	00264	02704	8	-0.00000	00000	00038	11696
9	0.00000	00000	00015	12912	9	-0.00000	00000	00000	13843
10	0.00000	00000	00000	94488	10	0.00000	00000	00000	00871
11	0.00000	00000	00000	02703	11	0.00000	00000	00000	00232
12	0.00000	00000	00000	00138	12	0.00000	00000	00000	00014
13	0.00000	00000	00000	00003					

Coefficients in the Expansion of

$$K_\nu(z) = (\pi/2z)^{\frac{1}{2}} e^{-z} \sum_{k=0}^{\infty} C_k(\nu) T_k^*(5/z), \quad z \geq 5, \quad ,$$

$$C_k(\nu) = \sum_{r=0}^{\infty} L_{r,k} \Pi_r^*(\nu), \quad 0 \leq \nu \leq 1. \quad .$$

r	$L_{r,k}, k = 6$			
0	-0.00000	00002	37980	67584
1	-0.00000	00047	09153	37960
2	-0.00000	00002	12579	69017
3	0.00000	00005	00775	96094
4	0.00000	00000	24837	13998
5	-0.00000	00000	10973	23314
6	-0.00000	00000	00559	83984
7	0.00000	00000	00075	24545
8	0.00000	00000	00003	99813
9	-0.00000	00000	00000	12459
10	-0.00000	00000	00000	00773
11	-0.00000	00000	00000	00020
12	-0.00000	00000	00000	00001

r	$L_{r,k}, k = 7$			
0	0.00000	00000	19822	45732
1	0.00000	00004	47354	03547
2	0.00000	00000	17667	05974
3	-0.00000	00000	48496	56295
4	-0.00000	00000	02104	49873
5	0.00000	00000	01133	89565
6	0.00000	00000	00050	47884
7	-0.00000	00000	00009	13484
8	-0.00000	00000	00000	41875
9	0.00000	00000	00000	02557
10	0.00000	00000	00000	00126

r	$L_{r,k}, k = 8$			
0	-0.00000	00000	01880	59807
1	-0.00000	00000	47495	63125
2	-0.00000	00000	01673	32073
3	0.00000	00000	05221	69434
4	0.00000	00000	00202	16650
5	-0.00000	00000	00127	72511
6	-0.00000	00000	00005	06442
7	0.00000	00000	00001	13970
8	0.00000	00000	00000	04625
9	-0.00000	00000	00000	00409
10	-0.00000	00000	00000	00017

r	$L_{r,k}, k = 9$			
0	0.00000	00000	00198	60992
1	0.00000	00000	05533	54983
2	0.00000	00000	00176	49151
3	-0.00000	00000	00614	91767
4	-0.00000	00000	00021	55533
5	0.00000	00000	00015	55073
6	0.00000	00000	00000	55762
7	-0.00000	00000	00000	14899
8	-0.00000	00000	00000	00545
9	0.00000	00000	00000	00062
10	0.00000	00000	00000	00002

r	$L_{r,k}, k = 10$			
0	-0.00000	00000	00022	96222
1	-0.00000	00000	00697	91868
2	-0.00000	00000	00020	38398
3	0.00000	00000	00078	21592
4	0.00000	00000	00002	51096
5	-0.00000	00000	00002	02942
6	-0.00000	00000	00000	06659
7	0.00000	00000	00000	02049
8	0.00000	00000	00000	00068
9	-0.00000	00000	00000	00009

r	$L_{r,k}, k = 11$			
0	0.00000	00000	00002	86971
1	0.00000	00000	00094	30265
2	0.00000	00000	00002	54535
3	-0.00000	00000	00010	64136
4	-0.00000	00000	00000	31573
5	0.00000	00000	00000	28180
6	0.00000	00000	00000	00854
7	-0.00000	00000	00000	00296
8	-0.00000	00000	00000	00009
9	0.00000	00000	00000	00001

Coefficients in the Expansion of

$$K_{\nu}(z) = (\pi/2z)^{\frac{1}{2}} e^{-z} \sum_{k=0}^{\infty} C_k(\nu) T_k^*(5/z), \quad z \geq 5, \quad ,$$

$$C_k(\nu) = \sum_{r=0}^{\infty} L_{r,k} T_r^*(\nu), \quad 0 \leq \nu \leq 1. \quad .$$

r	$L_{r,k}, k = 12$			
0	-0.00000	00000	00000	38387
1	-0.00000	00000	00013	53879
2	-0.00000	00000	00000	34024
3	0.00000	00000	00001	53647
4	0.00000	00000	00000	04245
5	-0.00000	00000	00000	04137
6	-0.00000	00000	00000	00117
7	0.00000	00000	00000	00045
8	0.00000	00000	00000	00001

r	$L_{r,k}, k = 13$			
0	0.00000	00000	00000	05453
1	0.00000	00000	00002	05154
2	0.00000	00000	00000	04830
3	-0.00000	00000	00000	23394
4	-0.00000	00000	00000	00606
5	0.00000	00000	00000	00639
6	0.00000	00000	00000	00017
7	-0.00000	00000	00000	00007

r	$L_{r,k}, k = 14$			
0	-0.00000	00000	00000	00817
1	-0.00000	00000	00000	32632
2	-0.00000	00000	00000	00724
3	0.00000	00000	00000	03736
4	0.00000	00000	00000	00091
5	-0.00000	00000	00000	00103
6	-0.00000	00000	00000	00003
7	0.00000	00000	00000	00001

r	$L_{r,k}, k = 15$			
0	0.00000	00000	00000	00129
1	0.00000	00000	00000	05424
2	0.00000	00000	00000	00114
3	-0.00000	00000	00000	00623
4	-0.00000	00000	00000	00014
5	0.00000	00000	00000	00017

r	$L_{r,k}, k = 16$			
0	-0.00000	00000	00000	00021
1	-0.00000	00000	00000	00938
2	-0.00000	00000	00000	00019
3	0.00000	00000	00000	00108
4	0.00000	00000	00000	00002
5	-0.00000	00000	00000	00003

r	$L_{r,k}, k = 17$			
0	0.00000	00000	00000	00004
1	0.00000	00000	00000	00168
2	0.00000	00000	00000	00003
3	-0.00000	00000	00000	00019
4	-0.00000	00000	00000	00000
5	0.00000	00000	00000	00001

r	$L_{r,k}, k = 18$			
0	-0.00000	00000	00000	00001
1	-0.00000	00000	00000	00031
2	-0.00000	00000	00000	00001
3	0.00000	00000	00000	00004

r	$L_{r,k}, k = 19$			
0	0.00000	00000	00000	00000
1	0.00000	00000	00000	00006
2	0.00000	00000	00000	00000
3	-0.00000	00000	00000	00001

r	$L_{r,k}, k = 20$			
0	-0.00000	00000	00000	00000
1	-0.00000	00000	00000	00001

$$(18) \quad d_{r,k} = (2/n) \sum_{\alpha=0}^{n-1} C_k(\mu_\alpha, \lambda) \cos r\theta_\alpha, \quad \mu_\alpha = \frac{1}{2}(1 + \cos \theta_\alpha),$$

$$(19) \quad e_{r,k} = (2/n) \left[\frac{1}{2} \{ C_k(1, \lambda) + (-1)^r C_k(0, \lambda) \} + \sum_{\alpha=1}^{n-1} C_k(\eta_\alpha, \lambda) \cos r\varphi_\alpha \right],$$

$$\eta_\alpha = \frac{1}{2}(1 + \cos \varphi_\alpha),$$

where θ_α and φ_α are defined by (10) and (14), respectively. Then omitting remainder terms, we have

$$(20) \quad L_{r,k}(\lambda) = \frac{1}{2}(d_{r,k} + e_{r,k}), \quad k = 0, 1, \dots, n - 1.$$

Finally, it is of interest to record a useful asymptotic estimate of $L_{r,k}(\lambda)$ for k sufficiently large. Since $T_r^*(\nu) = T_r(2\nu - 1)$, from (7) and (17),

$$(21) \quad L_{r,k}(\lambda) = \pi^{-1} \int_0^1 [\nu(1 - \nu)]^{-1/2} C_k(\nu, \lambda) T_r^*(\nu) \, d\nu.$$

If λ is fixed, we can insert (6) in (21), as the $O(k^{-1/3})$ term in (6) is uniform for ν such that $0 \leq \nu \leq 1$. Now,

$$(22) \quad \Gamma(\frac{1}{2} + \nu)\Gamma(\frac{1}{2} - \nu) = \pi \sec \nu\pi.$$

So

$$(23) \quad L_{r,k}(\lambda) = \beta_k A_r + \delta_k,$$

$$\beta_k = 4(-1)^k (\pi/3)^{1/2} k^{-2/3} (2\lambda)^{1/6}, \quad \delta_k = \beta_k O(k^{-1/3}),$$

$$A_r = \pi^{-1} \int_0^1 [\nu(1 - \nu)]^{1/2} T_r^*(\nu) \cos \nu\pi \, d\nu.$$

It can be readily shown that

$$(24) \quad A_r = 0 \quad \text{if } r \text{ is even,}$$

$$A_r = (-1)^{r+1} J_{2r+1}(\pi/2) \quad \text{if } r \text{ is odd,}$$

where $J_\mu(z)$ is the Bessel function of the first kind. Now,

$$(25) \quad J_\mu(z) = \frac{(z/2)^\mu}{\Gamma(\mu + 1)} [1 + O(\mu^{-1})].$$

Thus, we can easily estimate the number of coefficients needed in our double Chebyshev series expansions to achieve a given level of accuracy.

3. Numerical Results. We have from (1) and (17), with a slight change of notation,

$$(26) \quad K_r(z) = (\pi/2z)^{1/2} e^{-z} \sum_{k=0}^{\infty} C_k(\nu) T_k^*(5/z), \quad z \geq 5,$$

$$C_k(\nu) = \sum_{r=0}^{\infty} L_{r,k} T_r^*(\nu), \quad 0 \leq \nu \leq 1.$$

In the tables (pp. 326–328), we present the coefficients $L_{r,k}$ which were computed by the techniques previously enunciated.

In the computations for $d_{r,k}$ and $e_{r,k}$, we chose n an even integer, actually $n = 20$. Then, as a check on the coefficients, we computed $C_k(\nu)$ with the aid of the values

of $d_{r,k}$ for $\nu = 0, 1/4, 1/3, 1/2, 2/3, 3/4$ and 1, and $C_k(\nu)$ with the aid of the values of $e_{r,k}$ for $\nu = 1/4, 1/3, 2/3$ and $3/4$. Note that none of these ν values were used to produce the respective coefficients, in view of (10) and (14). Since

$$(27) \quad (2z/\pi)^{1/2} e^z K_{1/2}(z) = 1, \quad \text{all } z,$$

$$(28) \quad C_0(\frac{1}{2}, \lambda) = 1, \quad C_k(\frac{1}{2}, \lambda) = 0, \quad k > 0,$$

and as

$$(29) \quad T_{2m}^*(\frac{1}{2}) = (-1)^m, \quad T_{2m+1}^*(\frac{1}{2}) = 0, \quad m = 0, 1, \dots,$$

we have

$$(30) \quad \sum_{r=0}^{\infty} (-1)^r d_{2r,k} = 1, \quad \text{if } k = 0, \nu = \frac{1}{2}, \\ = 0 \quad \text{if } k > 0, \nu = \frac{1}{2}.$$

The computations were designed to make the coefficients accurate to about 23S. In view of their application in (26), it is convenient to round the coefficients to 20D for presentation. Then only 205 coefficients are needed to produce values of $(2z/\pi)^{1/2} e^z K_\nu(z)$ for all $z \geq 5$ and all $\nu, 0 \leq \nu \leq 1$, to 20D except possibly for round off. Actually, we have in effect coefficients for the evaluation of the above transcendental for all $z \geq 5$ and all $\nu \geq 0$, since $K_\nu(z) = K_{-\nu}(z)$ and use of the recurrence formula for $K_\nu(z)$, namely,

$$(31) \quad K_{\nu+1}(z) = K_{\nu-1}(z) + (2\nu/z)K_\nu(z)$$

is stable in the forward direction. The space required for these coefficients to compute $K_\nu(z)$ for all ν (and $z \geq 5$) in comparison with that required for existing conventional tables for only a few values of ν manifests the miniaturization described.

4. Concluding Remarks. To the best of our knowledge, except for some coefficients reported by Clenshaw and Picken [5], use of double Chebyshev series has not received much attention. These authors give coefficients to 6D for the evaluation of the Bessel functions $J_\nu(z)$ and $I_\nu(z)$, based on their ascending series representations valid for $0 < z \leq 8$ and $-1 \leq \nu \leq 1$. As previously remarked, we are extending our work to cover all the Bessel functions.

5. Acknowledgement. I am indebted to Miss Rosemary Moran for her assistance in the development of the numerics.

Midwest Research Institute
425 Volker Boulevard
Kansas City, Missouri 64110

1. Y. L. LUKE, *The Special Functions and Their Approximations*. Vols. 1, 2, Mathematics in Science and Engineering, vol. 53, Academic Press, New York, 1969. See especially Vol. 1, p. 213; Vol. 2, pp. 25-28. MR 39 #3039; MR 40 #2909.

2. *IBID.*, Vol. 2, p. 26.

3. *IBID.*, Vol. 1, pp. 308-314.

4. *IBID.*, Vol. 2, pp. 339, 341, 360, 362, 365, 367.

5. C. W. CLENSHAW & S. M. PICKEN, *Chebyshev Series for Bessel Functions of Fractional Order*, National Physical Laboratory Mathematical Tables, vol. 8, Her Majesty's Stationery Office, London, 1966. MR 34 #2948.