

HERMITIAN SPACES IN GEODESIC CORRESPONDENCE¹

W. J. WESTLAKE

1. Coburn [1] has studied the problem of Hermitian spaces in geodesic correspondence. He found a necessary and sufficient condition for two Kähler spaces to be in geodesic correspondence and showed that such correspondence was impossible between a Kähler space and a Hermitian space. The problem of geodesic correspondence between two Hermitian spaces he left unsolved.

It will be shown that Coburn's first result is incorrect, as was pointed out by Bochner [2]; and we shall solve the general problem, giving necessary and sufficient conditions for two Hermitian spaces to be in geodesic correspondence.

2. The linear connection of a Hermitian space, H_n , is defined to be

$$\Gamma_{\beta\gamma}^{\alpha} = g^{\alpha\lambda*} \partial_{\gamma} g_{\beta\lambda*}.$$

We use the usual convention of Hermitian geometry that Greek indices run from 1 to n and Latin indices from 1 to $2n$ where $\alpha^* = \alpha + n$.

The torsion tensor S_{jk}^i is defined to be

$$S_{jk}^i = \Gamma_{[jk]}^i$$

and its vanishing characterises a Kähler space, K_n . We shall also introduce a connection E_{jk}^i which is analogous to the Christoffel symbol

$$\left\{ \begin{matrix} i \\ jk \end{matrix} \right\}$$

of Riemannian geometry. This is:

$$E_{jk}^i = \frac{1}{2} g^{ih} (\partial_k g_{jh} + \partial_j g_{kh} - \partial_h g_{jk})$$

and may also be written in the form:

$$(2.1) \quad E_{jk}^i = \Gamma_{(jk)}^i + g^{ip} g_{jq} S_{pk}^q + g^{ip} g_{kq} S_{pj}^q.$$

For a Kähler space, it is clear that E_{jk}^i reduces to Γ_{jk}^i .

Received by the editors June 30, 1953.

¹ This paper is a short extract from a thesis approved by the University of London for the degree of Ph.D. It was written while the author was the recipient of a D.S.I.R. research grant.

The differential equations of the geodesics of a Hermitian space (given by Coburn [1]) may be rewritten in terms of E_{jk}^i and are:

$$(2.2) \quad \frac{d^2z^i}{dt^2} + E_{jk}^i \frac{dz^j}{dt} \frac{dz^k}{dt} - \frac{dz^i}{dt} \left(\frac{d^2s}{dt^2} \bigg/ \frac{ds}{dt} \right) = 0$$

with t as parameter. Multiplying by dz^m/dt , interchanging i and m , and subtracting, we have the alternative form:

$$(2.3) \quad \left(\frac{d^2z^i}{dt^2} \frac{dz^m}{dt} - \frac{d^2z^m}{dt^2} \frac{dz^i}{dt} \right) + \left(E_{jk}^i \frac{dz^m}{dt} - E_{jk}^m \frac{dz^i}{dt} \right) \frac{dz^j}{dt} \frac{dz^k}{dt} = 0.$$

3. Consider two Hermitian spaces H_n, H'_n whose geodesics correspond. The respective differential equations of their geodesics will be satisfied by the same functions $z^i(t)$ and subtracting we shall have:

$$(3.1) \quad (A_{jk\delta_n}^i \delta_n^m - A_{jk\delta_n}^m \delta_n^i) \frac{dz^n}{dt} \frac{dz^j}{dt} \frac{dz^k}{dt} = 0$$

where the symmetric tensor A_{jk}^i is defined by

$$(3.2) \quad A_{jk}^i = E_{jk}^{\prime i} - E_{jk}^i$$

and primed expressions refer to H'_n . Equation (3.2) must be satisfied by arbitrary values of $d_t z^i$ at any point. Therefore we have:

$$(3.3) \quad A_{(jk\delta_n)}^i \delta_n^m - A_{(jk\delta_n)}^m \delta_n^i = 0,$$

and by virtue of the symmetry of A_{jk}^i , this may be written:

$$(3.4) \quad A_{jk\delta_n}^i \delta_n^m + A_{kn\delta_j}^i \delta_j^m + A_{nj\delta_k}^i \delta_k^m = A_{jk\delta_n}^m \delta_n^i + A_{kn\delta_j}^m \delta_j^i + A_{nj\delta_k}^m \delta_k^i.$$

4. If both spaces are Kähler, A_{jk}^i has all its components identically zero except those of the type $A_{\beta\gamma}^\alpha, A_{\beta^*\gamma^*}^{\alpha^*}$. Putting $i = \alpha, j = \beta, k = \gamma, m = \mu^*, n = \nu^*$ in (3.4), we see that

$$(4.1) \quad A_{\beta\gamma}^\alpha = 0.$$

And conversely if (4.1) holds, (3.1) is satisfied identically. But

$$A_{\beta\gamma}^\alpha = \Gamma_{\beta\gamma}^{\prime\alpha} - \Gamma_{\beta\gamma}^\alpha$$

and thus:

A necessary and sufficient condition for two Kähler spaces, K_n, K'_n , to be in geodesic correspondence is that

$$\Gamma_{\beta\gamma}^{\prime\alpha} = \Gamma_{\beta\gamma}^\alpha.$$

This is the correction of Coburn's result referred to.

5. Suppose now that both spaces are Hermitian. Then equation (3.4) must again be satisfied. Putting $i = \alpha, j = \beta, k = \gamma, m = \mu^*, n = \nu^*$, we have

$$(5.1) \quad A_{\beta\gamma}\delta_{\nu^*}^{\mu^*} = A_{\gamma\nu^*}\delta_{\beta}^{\alpha} + A_{\mu^*\beta}\delta_{\gamma}^{\alpha} = 0.$$

The only components of A_{jk}^i that now vanish identically are those of the type $A_{\beta^*\gamma^*}^{\alpha}, A_{\beta\gamma}^{\alpha^*}$. In particular, $A_{\gamma\nu^*}^{\mu^*}$ is a tensor; and the contraction $A_{\gamma\mu^*}^{\mu^*}$, which we denote by A_{γ} , is a vector. Contracting on μ^*, ν^* in (5.1), we have:

$$(5.2) \quad nA_{\beta\gamma}^{\alpha} = A_{\gamma}\delta_{\beta}^{\alpha} + A_{\beta}\delta_{\gamma}^{\alpha}.$$

This constitutes our first necessary condition.

In (3.4), putting $i = \alpha, j = \beta^*, k = \gamma, m = \mu, n = \nu$, and contracting on μ, ν we have:

$$(5.3) \quad nA_{\beta^*\gamma}^{\alpha} = A_{\beta^*}\delta_{\gamma}^{\alpha},$$

and this is our second necessary condition.

If now we assume that both these necessary conditions are satisfied, it is easy to verify that (3.4) is then identically satisfied for every possible combination of the indices, i, j, k, m, n . Thus (3.1) is identically satisfied and the two necessary conditions are also sufficient. Simplifying (5.2) and (5.3) and denoting by S_{α} the contraction $S_{\alpha\mu}^{\mu}$, our final result is:

The necessary and sufficient conditions for two Hermitian spaces H_n, H'_n to be in geodesic correspondence are:

$$n\Gamma'_{(\beta\gamma)}^{\alpha} + \delta_{(\beta}^{\alpha}S'_{\gamma)} = n\Gamma_{(\beta\gamma)}^{\alpha} + \delta_{(\beta}^{\alpha}S_{\gamma)},$$

$$ng'^{\alpha\lambda}g'_{\gamma\nu^*}S_{\lambda^*\beta^*}^{\nu^*} + \delta_{\gamma}^{\alpha}S'_{\beta^*} = ng^{\alpha\lambda}g_{\gamma\nu^*}S_{\lambda^*\beta^*}^{\nu^*} + \delta_{\gamma}^{\alpha}S_{\beta^*}.$$

BIBLIOGRAPHY

1. N. Coburn, *Unitary spaces with corresponding geodesics*, Bull. Amer. Math. Soc. vol. 47 (1941) pp. 901-910.
2. S. Bochner, *Curvature in Hermitian metric*, Bull. Amer. Math. Soc. vol. 53 (1947) pp. 179-195.